

# **Direct detection of dark sector DM via electron counting in liquid xenon**

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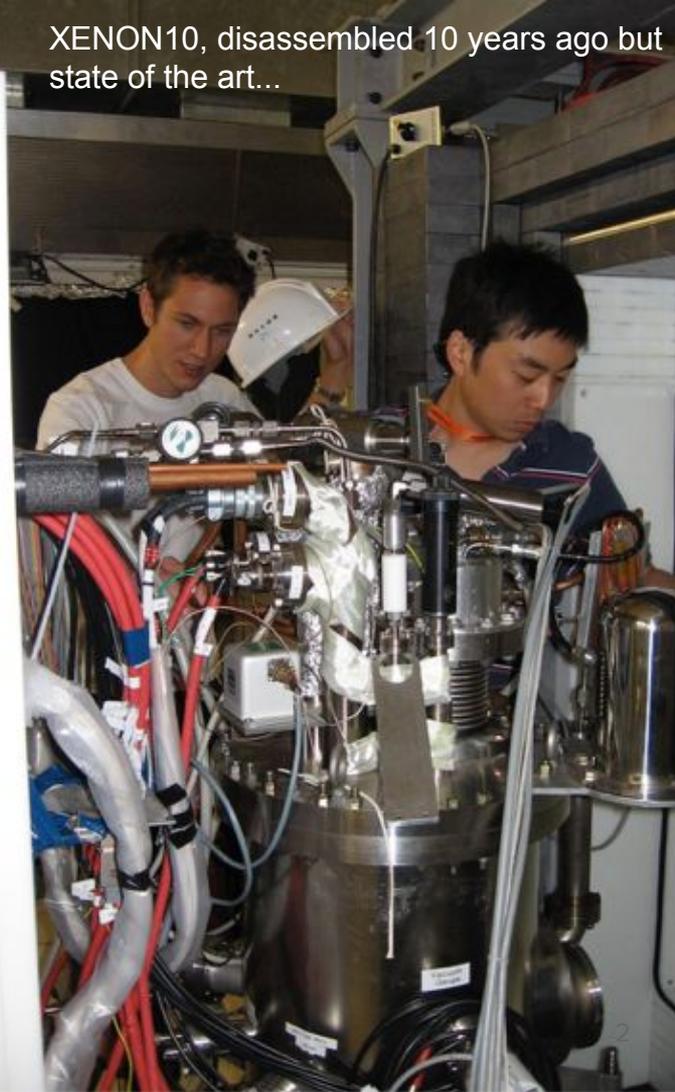
Peter Sorensen on behalf of the

**$U_A(1)$  Collaboration**

U.S. Cosmic Visions Workshop, 23-25 March 2017, College Park, Maryland

# $U_A(1)$ concept

- 10 kg scale liquid xenon TPC with complete focus on S2 signal and mitigation of e- backgrounds
- Without concern for S1 (primary scintillation collection)
  - the design is far simpler
  - and cheaper
  - contains less plastics (easier to achieve purity)
- A 2 kg scale prototype is already built
  - LLNL detector for CENNS
  - Update prototype design for 10 kg active while studying e- background mitigation
- Underground deployment at SURF
  - Small footprint, likely compatible with BLBF space



# Target mass versus atomic bandgap

- Xe has a large  $>9.2$  eV band gap, which suppresses the scattering rate
  - Semiconductors have  $\sim 1$  eV band gap, a distinct advantage, however...
- **Mass is a relentless advantage in direct detection**
- And, tonne-scale liquid xenon TPCs are being deployed and/or built
  - 1000+ kg xenon vs  $<1$  kg for semiconductors
  - It would be great to leverage large, quiet, sensitive targets (e.g. LZ) which are being deployed **anyway** for related purposes
- Even a 10 kg target can search new parameter space in the short term

# Sensitivity not guaranteed (unless!)

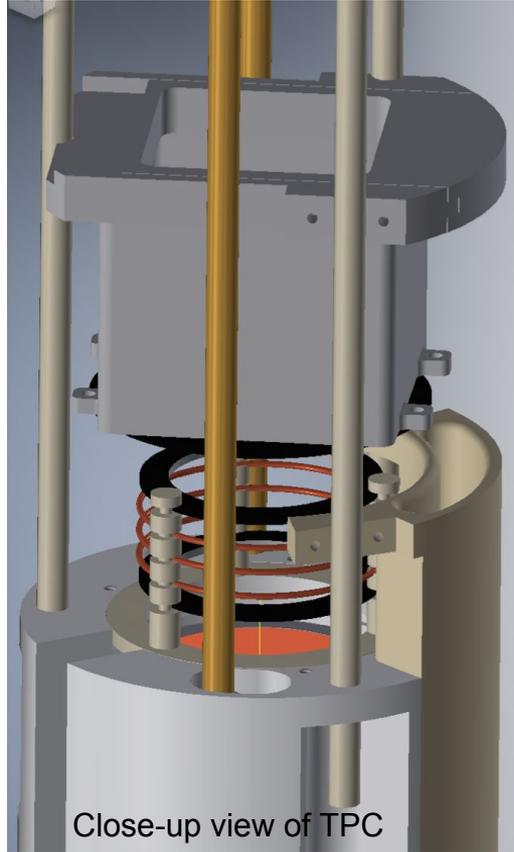
- Ability of LZ/XENON1T to do single electron analyses presently doubtful
  - XENON10: single electron sensitive search but limited by electron train background
  - XENON100: 4-5 electron threshold and still limited by background
  - LUX: in progress...
  - e- backgrounds have been considered a minor irritation to the primary goal of finding WIMPs
  - Efforts to mitigate them have so far been modest
- Mitigation requires a dedicated effort
  - Initial small-scale (surface) efforts underway (LLNL, LBL)
  - **Underground test bed eventually essential due to long lifetime of correlated backgrounds**
- Might as well get a science result in the process!

- ■ The  $U_A(1)$  experiment

# Prototype already built at LLNL



Super portable for “drop-ship” deployment



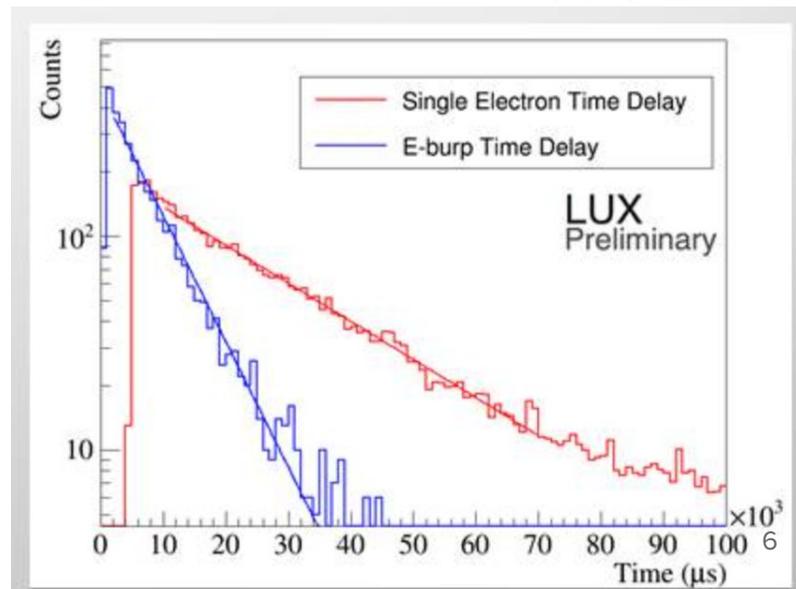
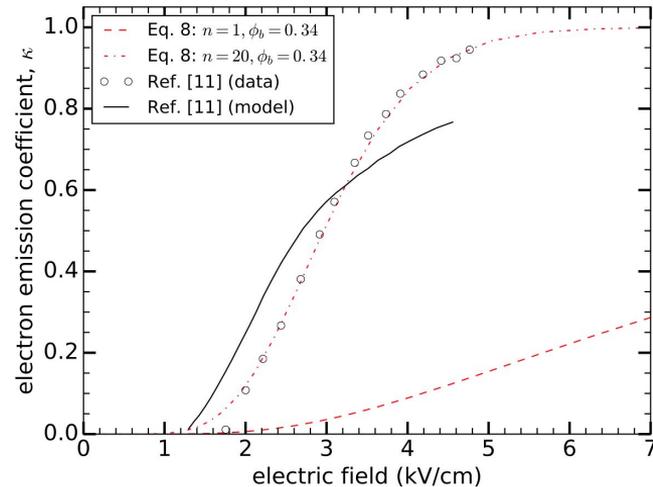
Close-up view of TPC



Full view of TPC

# Primary R&D is to control e- backgrounds

- LUX studies underway (Jingke Xu, LLNL)
  - e.g. talk at APS 2016 April meeting
  - Two primary classes of electron backgrounds
    - Single e- backgrounds
    - e- clusters
      - events tend to be quite large
      - So less of a concern for few e- counting
- Recent theoretical work on understanding thermal e- trapping (Sorensen, LBL)
  - Predicts trapping lifetime  $O(10)$  ms
  - arXiv:1702:04805
- Additional R&D is underway at LBL and LLNL

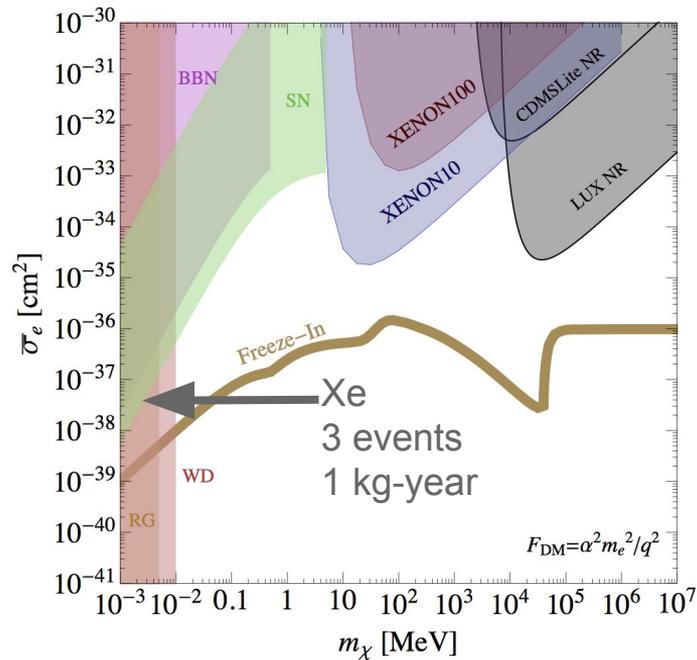
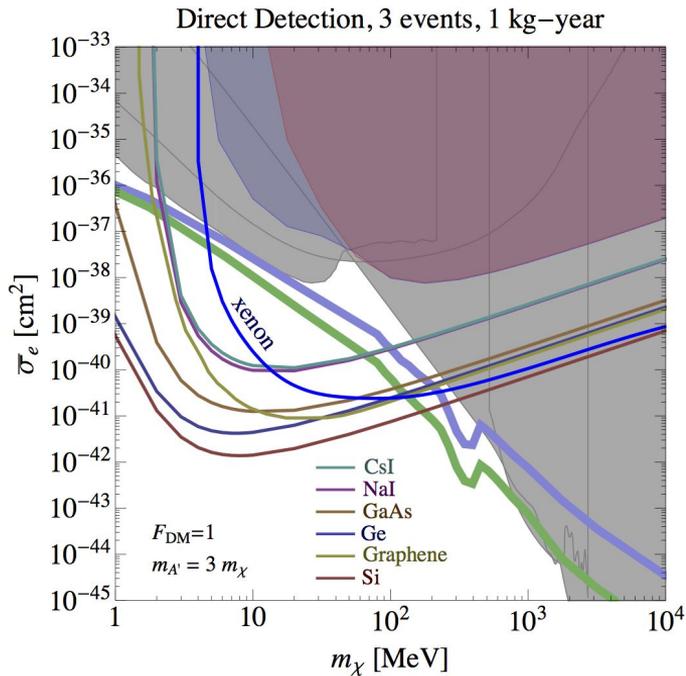


# Sources of electron backgrounds

Sources	Mitigations
Trapped electrons at the liquid gas interface	<ol style="list-style-type: none"><li>1. larger electron emission field</li><li>2. Infrared photons to liberate trapped e-</li><li>3. Last resort: HV switching</li></ol>
Spontaneous emission from metal surfaces <ol style="list-style-type: none"><li>A. Due to inhomogeneities</li><li>B. Due to lowered work function resulting from trapped ions</li></ol>	Varies... <ol style="list-style-type: none"><li>A. Treatment of metal surfaces</li><li>B. AC field to de-trap ions</li></ol>

# Reach thermal DM production parameter space in <1 year!

The only existing limits on dark sector DM are from liquid xenon targets



plots from Essig et al, cf. arXiv:1703:00910

# Timescale and budget

- One year to update design (based on LLNL prototype)
- One year to build and deploy at SURF
- We are talking about a 10 kg scale experiment so these are realistic estimates
- 6 months to commission and verify the success of the background mitigation strategies
- 6 months to obtain first results
- **3 years total project**
  - Of which approx 2 years include R&D
- Estimate \$3M project

# Summary

- Deploy a small O(10) kg liquid xenon TPC with a focus on electron counting and mitigation of e- backgrounds
  - A cost-effective fishing expedition with a clear target! (cf. Weiner talk, morning plenary)
- Potential for rapid exploration of new dark sector DM parameter space
  - Including freeze-out / freeze-in regions
  - Complementary to beam dump experiments
- Provide essential data on e- backgrounds such that much larger detectors can later also be sensitive to dark sector DM
- Leverage existing infrastructure, expertise and underground facility access within LUX/LZ/community
  - Interested in joining this effort? Contact Adam Bernstein and/or Peter Sorensen

# Additional details about mitigations

- larger electron emission field
  - XENON achieved  $\sim 5.5$  kV/cm
  - Suspect  $>7$  kV/cm needed for substantial reduction of e-train bkgd
- Infrared photons to liberate trapped e-
  - Liquid surface trapping potential is 0.34 eV
  - 940 nm LEDs readily available (1.3 eV photon), trigger on S2
- Last resort: HV switching
  - Divert trapped electrons back to gate electrode
  - Possible in principle, may actually work quite well

# Additional detail about dark counts

- From recent paper, arXiv:1702.04805
  - Xe liquid/gas interface presents a 0.34 eV potential barrier for e- dark counts
  - This gives a O(10) ms trapping lifetime

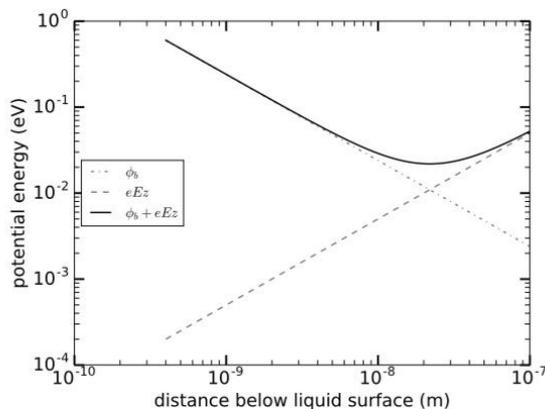


FIG. 2. Potential energy of an electron just below the liquid-gas interface.

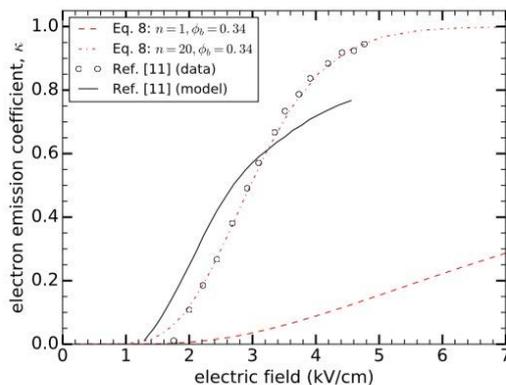


FIG. 1. Absolute efficiency for electron emission from liquid into gas xenon, as a function of the electric field in the liquid phase.

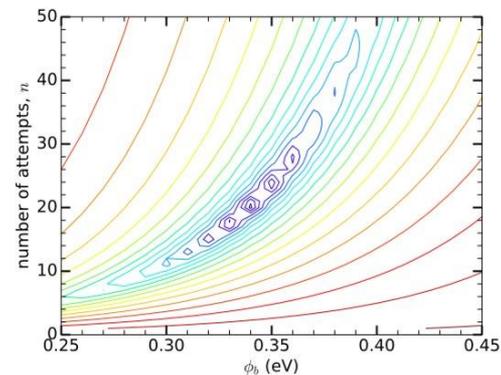
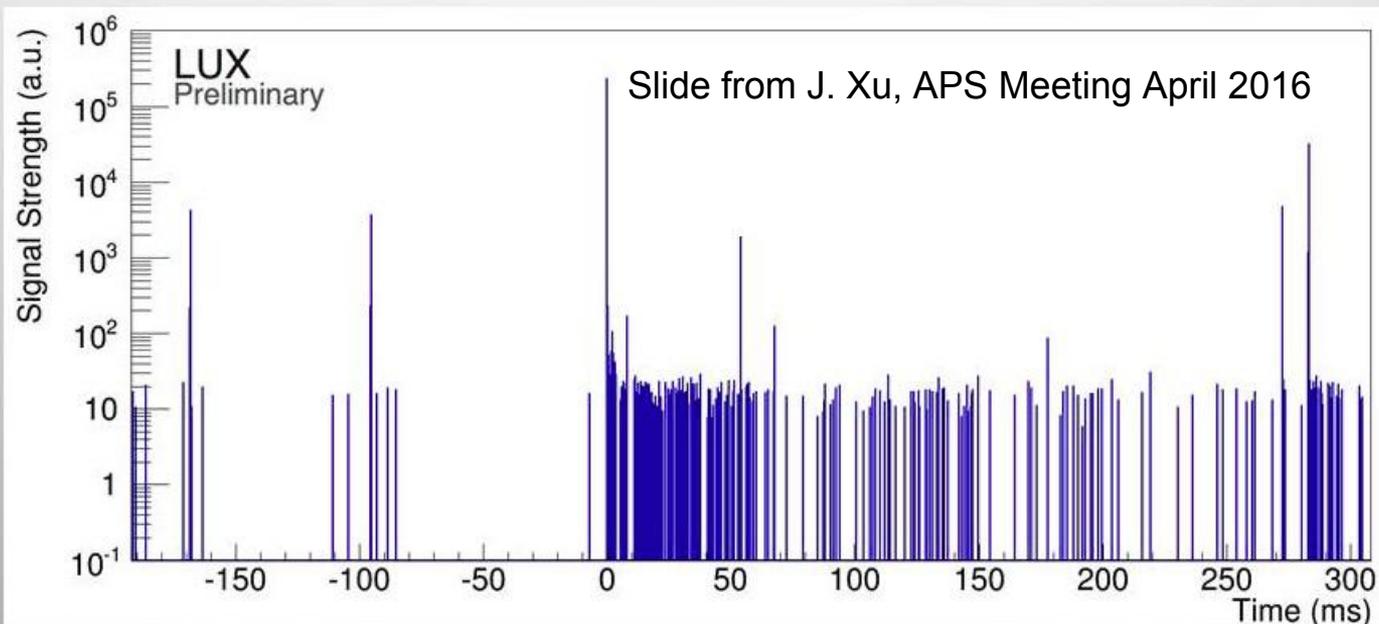


FIG. 4. Contour plot of  $\chi^2$  for comparison of the emission model given by Eq. 8 vs the data shown in Fig. 1. Agreement at  $2\sigma$  is confined to  $\phi_b = 0.34 \pm 0.1$  eV.

# Additional details about dark counts

A LUX event waveform over 500ms (maximum drift time in LUX: 0.325 ms)



- Temporal event correlations are observed at time scales 1000 times larger than typical LUX event.

# Additional plots

